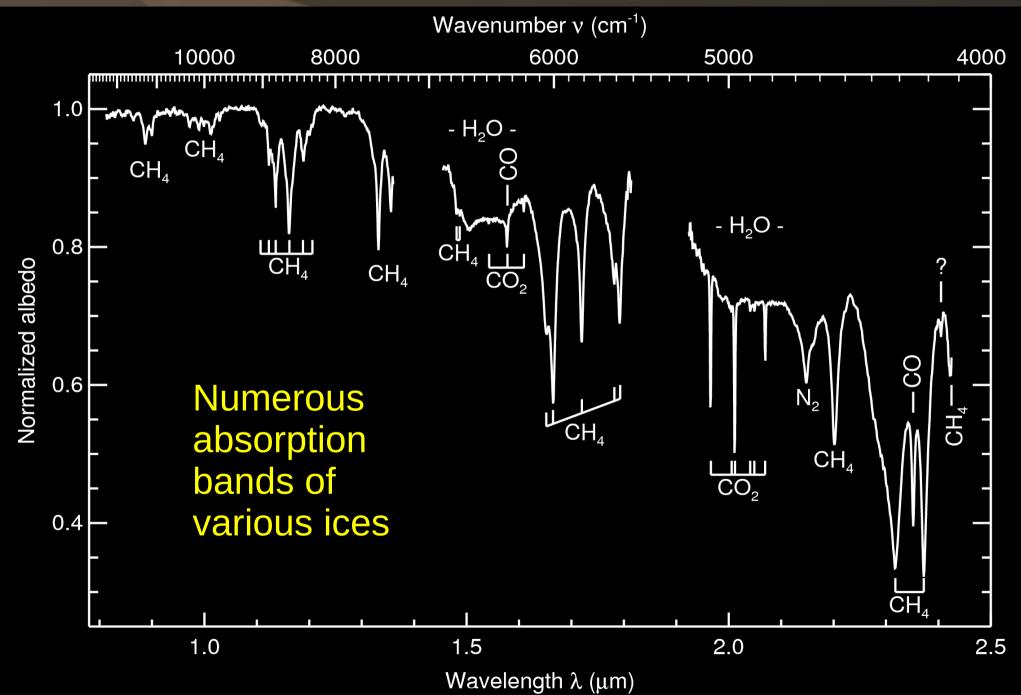
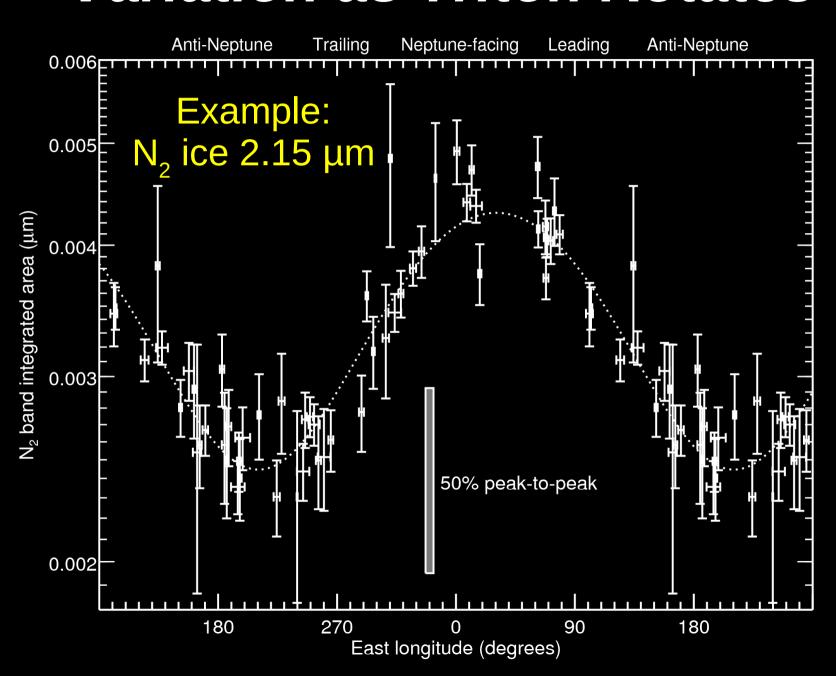
Ten Years of Triton Spectral Monitoring with IRTF/SpeX

- Observed every year since 2000, now more than 50 nights, looking for transient events. Takes just a few hours – perfect for remote observing.
- No sign of transient events so far, but spectacular detail on longitudinal distribution of ices, and also hints of secular evolution.
- Reference: Grundy, Young, Stansberry, Buie, Olkin, and Young 2010. Near-infrared spectral monitoring of Triton with IRTF/SpeX II: Spatial distribution and evolution of ices. Icarus 205, 594-604.

Average Triton Spectrum

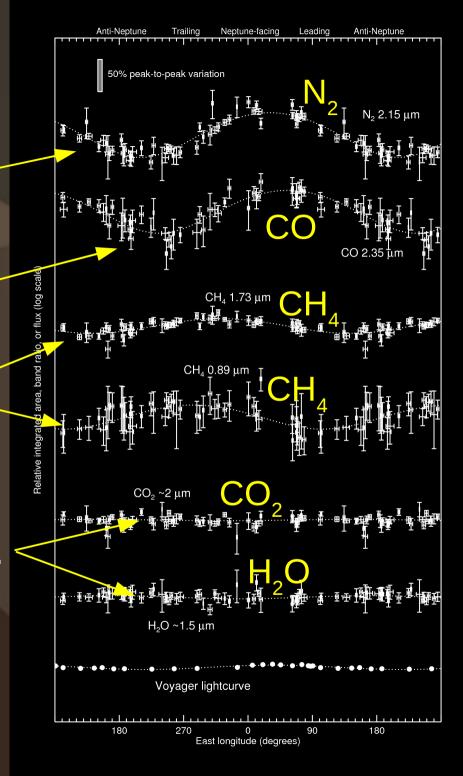


Measure Absorption Band Variation as Triton Rotates

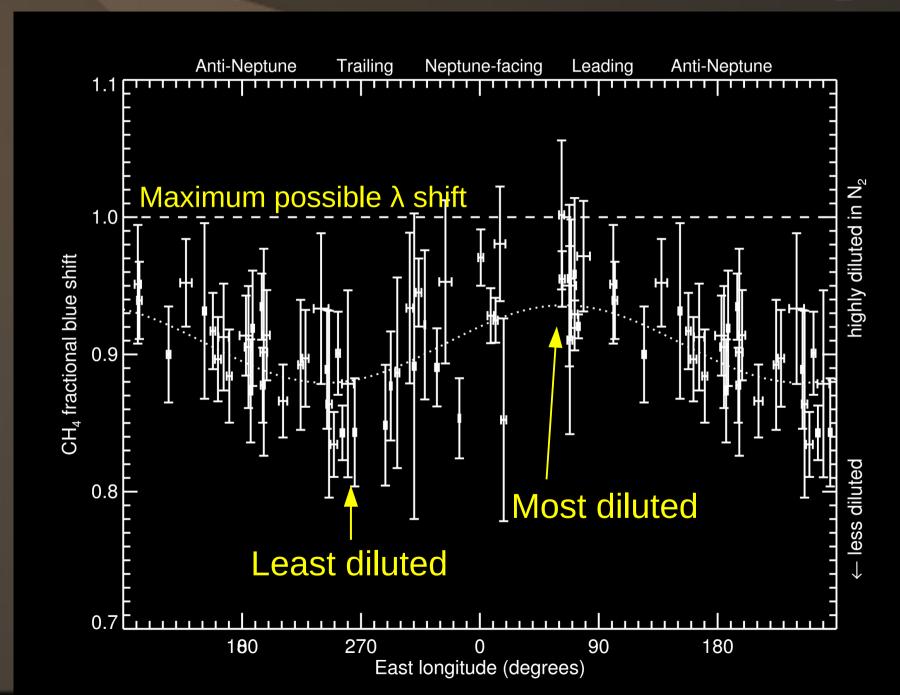


Triton Summary

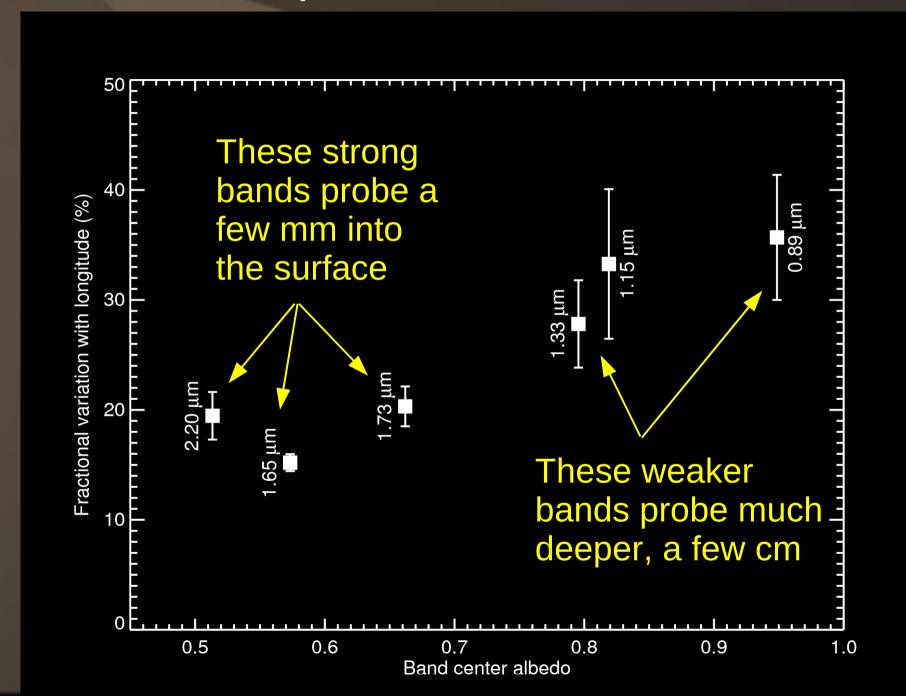
- Big N₂ ice enhancement on Neptune-facing hemisphere.
- CO ice is similar. Co-occurs with N₂ ice.
- CH₄ ice is different. More concentrated on trailing hemisphere.
- H₂O and CO₂ ices do not vary.
 Globally distributed, or only at high latitudes.



More CH_4 Details: Dilution in N_2 Ice



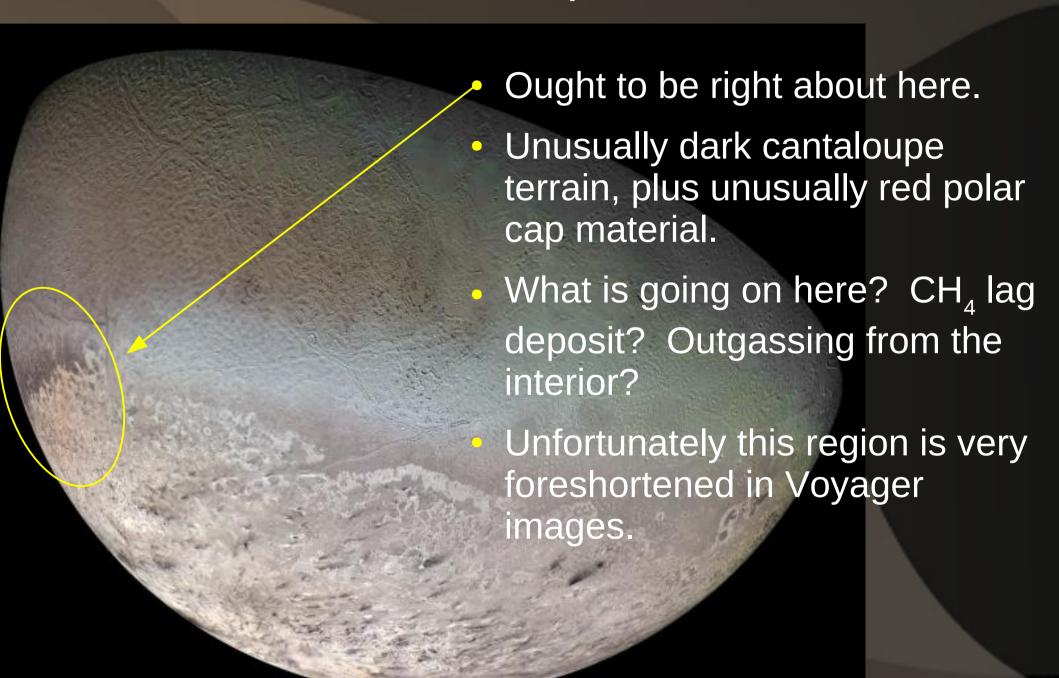
More CH, Details: Stratigraphy



Complex CH₄ Ice Behavior

- Longitudinal pattern is very different from N_2 ice, but wavelength shift says CH_4 is mostly diluted in N_2 ice.
- Wavelength shift varies with longitude, with least shift (least dilution) on the trailing hemisphere, where CH₄ absorption is strongest.
- Stronger CH₄ bands (which probe shallower) show lower amplitude variation than weaker bands (which probe deeper).
- Suggests a concentration of CH₄-rich ice below the surface, localized around 300° E longitude.

Where Is This CH₄-Rich Region?



Secular Evolution of N₂ Ice?

Dashed curve is sine fit to all data. Solid curves are sines fit to 2000-2004 and 2005-2009 subsets.

Amplitude drop is not quite 3-σ. Needs to be confirmed.

